# Whole Atmosphere Working Group Overview and Developments

CESM Joint AMWG, CCWG, ESPWG, CVCWG, and WAWG Winter WG Meeting

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11 June 2025



## WACCM & WACCM -X for CESM3

	Dyn. Core	Resolution	# levels	Chemistry	Physics
WACCM6	FV	1°, 2°	70 (110)	TSMLT, MA	CAM6
WACCM7	SE	ne30, ne16	135	t1ma, t4ma	CAM7
WACCM-X 2.1	FV	1°, 2°	130	TSMLT, MA, IT chemistry	CAM6
WACCM-X v?	SE	ne30, ne16 (ne120)	189 (273)	t1ma, t4ma IT chemistry	CAM7



#### Priority development objectives:

- Good QBO at both 1° and 2° resolutions
- Seasonal cycle of winds/temperatures in the stratosphere mesosphere and the impact on chemistry
- Thermosphere extension of SEdycore, ionospherethermosphere climatology



### WACCM7 / WACCM7 -X Status for CESM3

- QBO tuning is progressing, with QBO looking reasonable though not perfect
- Stratosphere, mesosphere, and lower thermosphere dynamics and seasonal cycle have been evaluated. No major issues, but still some room for improvement.
- Updated nitrogen chemistry in lower thermosphere which improves thermosphere temperature and neutral density in WACCNX
- Updating Shapiro filter in WACCMX for SEdycore
- Gravity wave parameterization tuning for highesolution (NE120) WACCMX
- ADF now supports more WACCMfocused results
- Known issues/biases: robust calculation of TEM diagnostics for Selycore, water vapor tape recorder in WACCM





Operated by UCAR

Dunkerton and Delisi QBO amplitude (ms<sup>-1</sup>)

## Results from base -runs – Polar Vortex (DJF)







## Results from base -runs - Polar Vortex (JJA)







## Results from base -runs - Seasonal Cycle (NH)



2 deg



## Results from base -runs – Seasonal Cycle (SH)







### **Current/Future WAWG Development Plans**

- WACCM
  - $\circ$   $\,$  New method to calculate zonal means for TEM fluxes  $\,$
  - In-line photolysis with TUV -x
  - WACCM MPAS development
- WACCM-X
  - Two-way coupling with GAMERA magnetosphere model
  - Improved ionosphere dynamo and transport
  - Scale aware gravity wave parameterization & thermosphere GW parameterization
  - Interface with JEDI for space weather data assimilation
- Reducing model cost through ML applications
  - Potential venues include ML troposphere, ML aerosol microphysics



## **Alternative Zonal Means and TEM diagnostics**

- Use ESMF library to remap dynamical quantities from unstructured physics grid to a distributed regular longitudelatitude grid
- Straight forward to compute zonal-means and TEM diagnostics on the regular longitude latitude grid
- Inline parallel ESMF regridding capabilities are efficient
- Adds only 1-2% to overall cost of WACCM
- Working on understanding potential noise issues at high latitudes



Francis V. Justin R.



## **In-line Photolysis**

- Photolysis was previously calculated by first using a lookup table (LUT)
- Using this approach photolysis rates did not include the direct effects on actinic flux from aerosol scattering and absorption.
- Now in-line calculation of the actinic flux.
- uses in-line Tropospheric ultraviolet-extended (TUV-x) allows the actual aerosol profiles in the column to be included in the calculation.
- For the aerosols, the optical properties (aerosol optical depth, single-scattering albedo, and asymmetry factor) calculated for the coarser band structure used by the radiative transfer code have been interpolated onto the wavelength grid used by TUV-x.

E.g., after Pinatubo eruption, the sulfate aerosols layer in the stratosphere can affect the photolysis rates in the atmosphere.



# WACCM – MPAS implementation

#### Previous development

- Stable WACCM/MPAS specified chemistry configuration for 120 km horizontal mesh with 70 vertical levels.
- <u>Development of WACCM with the non-hydrostatic MPAS-A dynamical core</u> S Kamali, HL Liu, W Skamarock, J Klemp, F Vitt... Journal of Advances in Modeling Earth Systems, 2024

#### Currently under development

- Stable configuration for 120 km horizontal mesh with 135 vertical levels for WACCM/MPAS with full chemistry with the goal of studying QBO.
- Constant pressure upper boundary development in MPAS standalone underway that requires more testing before bringing the changes into CESM. This is being developed specifically for geospace applications.
- Regional refinement to 3 km over CONUS (SIMA S2S^2)





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# WACCM-X/GAMERA: Towards a whole geospace model

- Forcing from the magnetosphere at higHatitudes is one of the largest uncertainties in simulating the upper atmosphere response to geomagnetic storms
- Empirical models fail to reliably capture the highatitude forcing
- NASA Center forGeospace Storms (CGS) and HAO have recently completed *two-way coupling* between WACCM-X and the GAMERA magnetosphere model, paving the way for a whole geospace model

#### Model Framework for CGS MAGE





MAGE: Multiscale Atmosphere-Geospace Environment Model



### Summary and Remaining Development Tasks

- Stratosphere, mesosphere, and lower thermosphere dynamics and seasonal cycle have been evaluated in baseline runs
- In-line photolysiscalculation with TUVx
- WACCM goes MPAS
- WACCM-X is heading towards wholegeospace model

#### **Development Tasks:**

- TEM calculation
- Low-frequency tropical GW parameterization
- WACCM with CAM7 93I vertical grid
- Explore ML options for tuning and replacing troposphere
- In-line photolysis calculation

